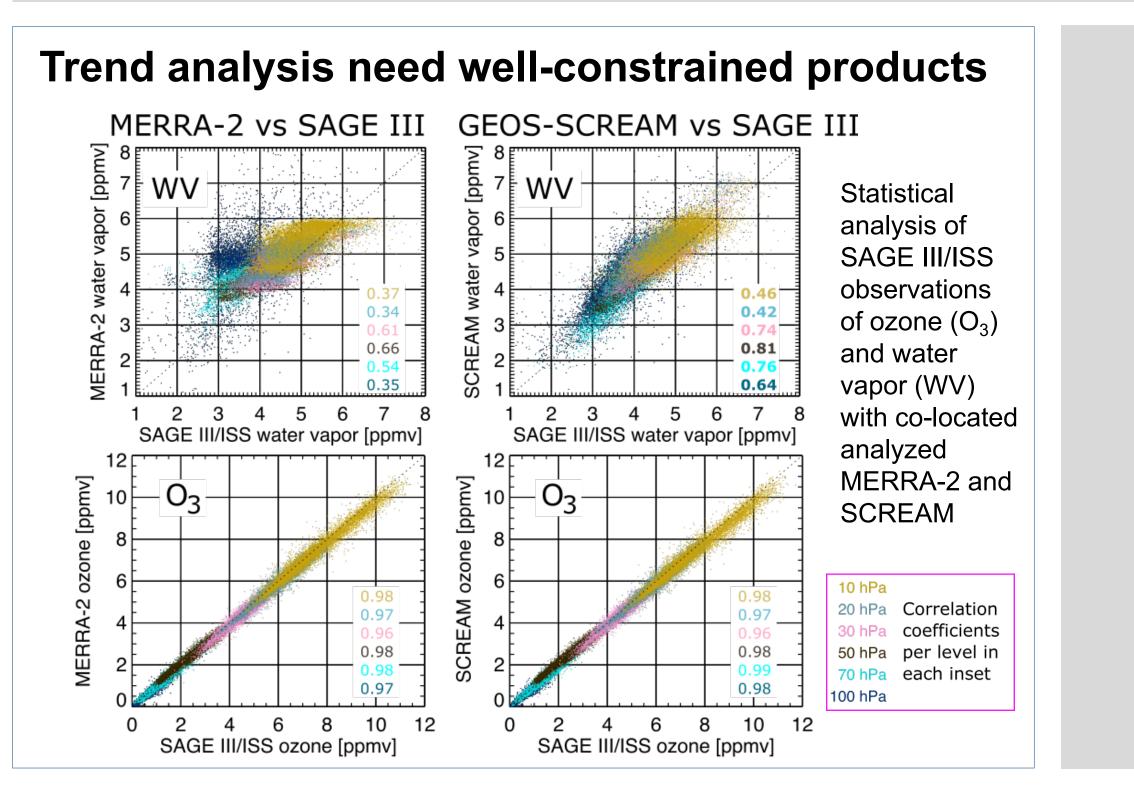
Bridging the SAGE data gap: Toward a climate data product with ozone and water vapor data from NASA SAGE and Aura missions and NASA reanalyses K. Emma Knowland^{1,2}, Steven Pawson², Pam Wales^{1,2}, Krzysztof Wargan^{2,3} and Brad Weir^{1,2}

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Main Goal for Project

Use NASA's uniformly-gridded, global GEOS model and data assimilation (DAS) products to bridge the gap between the earlier SAGE missions and the SAGE III/ISS products.

Question1: Can we use SAGE II ozone measurements to extend the ozone trend analysis across the 1998 observing system change?

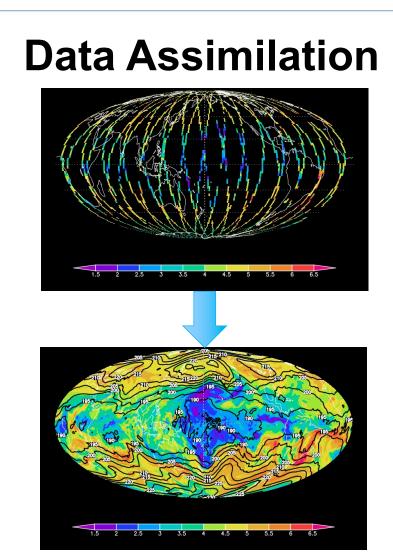


SAGE III/ISS & II Water vapor

- Annual Global Mean stratospheric Water Vapor (H₂O) from the SCREAM reanalysis agrees well with SAGE III/ISS for 2017 through 2020 (2018) shown ⊇).
- The vertical profiles for SAGE II are smoothed whereas they are not for SAGE III/ISS, evidenced by the larger standard deviation for SAGE III/ISS above 35 km

What is Data Assimilation

Bayesian method of combining and propagating information from observations in space and time using the governing equations and error estimates



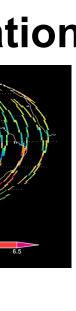
Example, 2 January 2016 at 100 hPa. **Top:** MLS water vapor. Bottom: Assimilated MLS water vapor (color) and MERRA-2 temperature (lines)

SAGE III/ISS vs GEOS Reanalyses Reanalysis stratospheric WV historically poor without observational constraint; correlation (r) improves with SCREAM vs SAGE III/ISS Despite differences in complexity of stratospheric O₃ chemistry, both reanalyses near perfect r with SAGE **SAGE III/ISS & SAGE II Water Vapor retrievals differ** SCREAM minus SAGE II Left figures characterize the difference between SCREAM co-located with SAGE II (2005) and SAGE III/ISS (2018). At each altitude bin is a pdf of "SCREAM minus SAGE" differences (color) and mean (circle) and median (cross) differences. The instrument precision (dotted lines) agree well with standard deviation (thick dashed lines) except at high altitudes and UTLS. The global mean concentrations of water vapor (right) for SCREAM is biased low at high altitudes compared to SAGE II while biased high throughout vs SAGE III/ISS. SCREAM H₂O minus SAGE III/ISS SCREAM minus MLS v5 00S 45S EQ 45N 90N EAM minus ACE-FTS **GEOS-SCREAM** assimilated MLS v4.2 water vapor which is biased high. We will need to 1) bias 90S 45S EQ 45N 90N correct the SCREAM water vapor in line with MLS v5 before we can use it in trend analysis bin size 0.20 frequency with SAGE water vapor observations and 2) keep in mind that the ACE-FTS measurements (used as

independent observations for validation) are also

biased in reference to MLS.

0.00 0.01 0.02 0.04 0.05 0.00



GEOS DAS and CCMM products

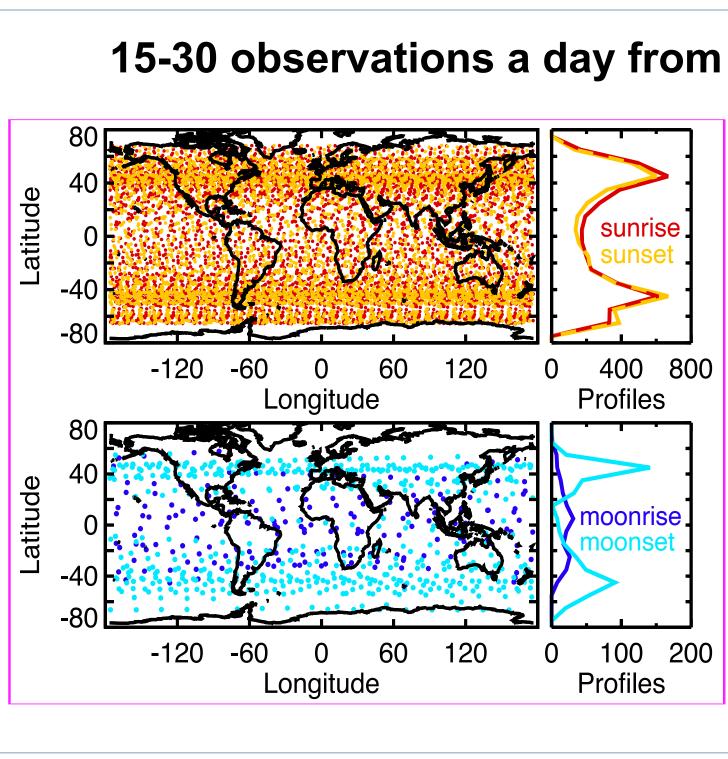
- **MERRA-2**: Meteorological reanalysis with assimilated ozone
- **GEOS-SCREAM**: Stratospheric Composition Reanalysis with Aura MLS using GEOS Constituent DAS "CoDAS" framework **MERRA2-GMI**: GEOS CCMM with GMI

Question 2: Can the assimilation of SAGE water vapor and ozone support trend and climate assessments after the Aura mission?

Constituent DA

Chemical data assimilation of O₃ and WV profiles

- > SAGE data is likely suitable for assimilation into GEOS using CoDAS framework
- Expectation water vapor will have more impact.



Sensitivity experiments

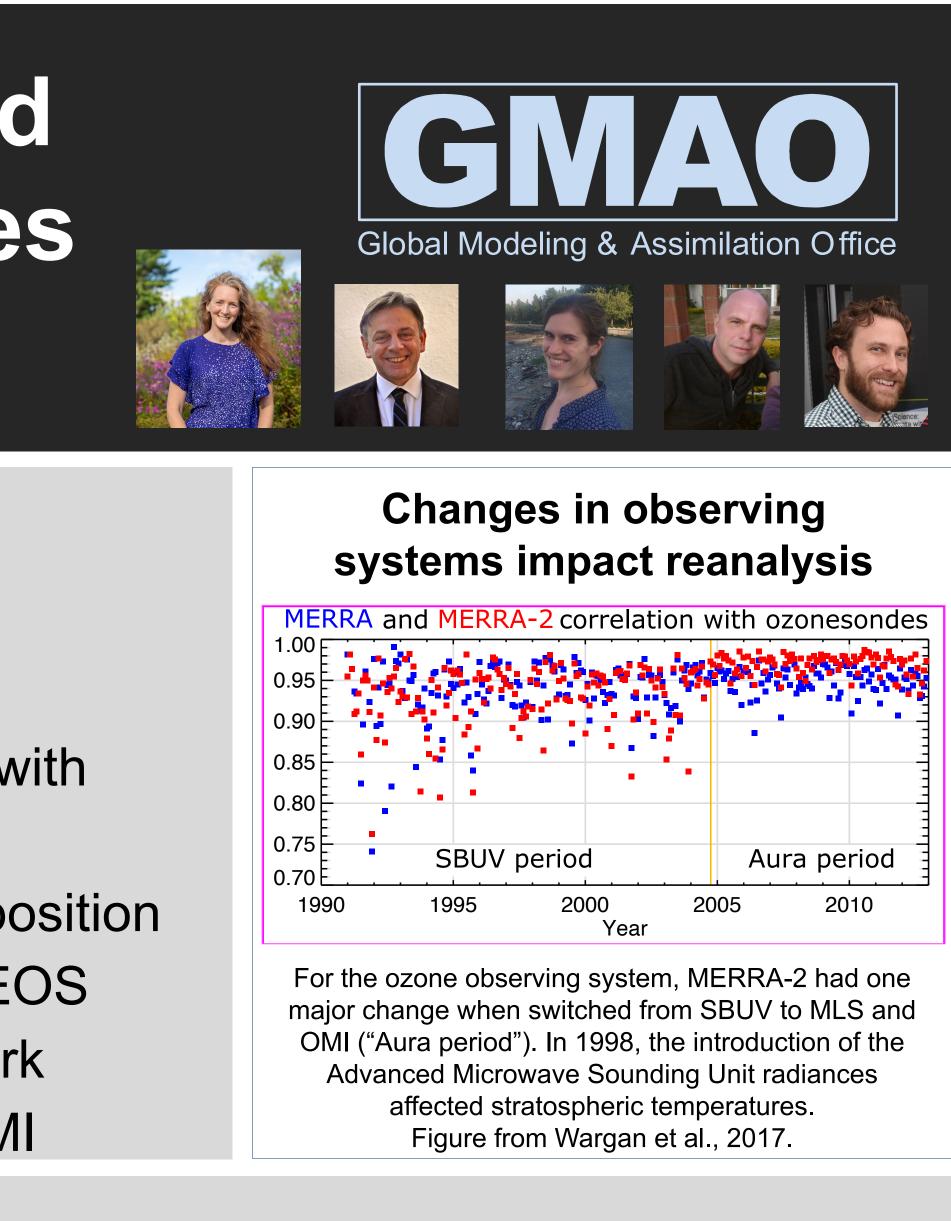
Same set up as GEOS SCREAM, with GEOS SCREAM as initial conditions. Evaluation will be done with ACE-FTS, frost-point hygrometers, sondes, lidar **Control**:

- 1. Coupled Chemistry and Meteorology Model (similar to MERRA2-GMI)
 - 1. 2016 simulation to test how CCMM may deviate from SCREAM (running)
 - 2. June 2017 to 2021+ control to coincide with CoDAS experiments

GEOS CoDAS:

- SAGE III/ISS + Aura (MLS and OMI) (testing begun with this configuration)
- 2. SAGE III/ISS + Aura alternative (e.g., OMPS)
- 3. SAGE III/ISS O₃ and WV assimilated only





15-30 observations a day from 70 °S to 70 °N

Ozone is available from both solar and lunar occultation measurements while water vapor is only available from the solar occultation.

Lower stratospheric O_3 and WV have chemical timescales long enough that 15-30 solar occultation observations a day can have a positive influence on the analyzed fields. We will use 3DVar assimilation with 6hour cycle windows, reducing the number of occultations to 3-8 per DAS cycle.



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