

Ensemble background-error variances: objective filtering and impact studies L. Raynaud, L. Berre and G. Desroziers CNRM/GAME, Météo-France/CNRS, Toulouse, France

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- \bullet Key element of any DA schemes : background-error covariance matrix ${\bf B}$
- ENDA provides a suitable framework to estimate **B**
 - Simulation of the estimation errors along analyses and forecasts
 - Documentation of error covariances :
 - over a long period
 - \Rightarrow "climatological error"
 - for a particular date \Rightarrow "error of the day"

(Evensen, 1997; Fisher, 2004; Berre et al., 2007)



.Introduction

2.Spatial structure of sampling noise 3.Design of an objective filter 4.Application to a NWP context 5.Impact studies 6.Conclusions and perspectives

1.1.ENsemble Data Assimilation 1.2.Main issues

- Only small size ensembles $(10 \rightarrow \mathcal{O}(10^2))$ are affordable \implies detrimental sampling noise for the estimation of **B**:
 - noisy variance fields (Berre et al., 2007; Raynaud et al., 2008)

- spurious non-zero correlations at long distances (Houtekamer and Mitchell, 1998; Buehner and Charron, 2007; Pannekoucke et al., 2007)

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- Present an application of the filter to a real NWP ensemble
- Estimate the impact of this filter on forecast scores
- Give some more general results about the benefits of using errors of "the day"

2.1.Empirical insight 2.2.Analytical results



(a)
$$N = 50, L_{\epsilon^b} = 200 \text{km}$$



Spatial structure of sampling noise (Fisher and Courtier 1995 (Fig 6), Raynaud et al., 2008)

 $\begin{array}{l} \text{True variance field} \\ \mathbf{V}^{\star} \sim \text{large scale} \\ & \downarrow \\ \text{Sampling noise} \\ \mathbf{V}^{e} = \tilde{\mathbf{V}}(N) - \mathbf{V}^{\star} \sim \text{large scale too} \, ? \\ & \rightarrow \text{ depend on } L_{\epsilon^{b}} \end{array}$

 $\implies Close link between the spatial structures of sampling noise and background-error$

2.1.Empirical insight 2.2.Analytical results

Notations

- $-\tilde{\mathbf{B}}$: the estimated \mathbf{B} matrix,
- $-\tilde{\mathbf{B}}^{\star} = E[\tilde{\mathbf{B}}]$: the noise-free estimated **B** matrix,
- $-\mathbf{V}^e = \tilde{\mathbf{B}} \tilde{\mathbf{B}}^{\star}$: the sampling noise or random error component.

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- $-\mathbf{V}^e = \tilde{\mathbf{B}} \tilde{\mathbf{B}}^{\star}$: the sampling noise or random error component.
- Analytically, it can be shown that the noise covariance matrix is

$$E[\mathbf{V}^e \mathbf{V}^{e^T}] = \frac{2}{N-1} \tilde{\mathbf{B}}^\star \circ \tilde{\mathbf{B}}^\star$$

where \circ stands for the Hadamard product :

- spatial structures of sampling noise and background-error are directly related,
- the relative error of the variance estimation, $\frac{E[(\mathbf{V}^e)^2]}{(\mathbf{V}^\star)^2} = \frac{2}{N-1}$, is inversely proportional to the ensemble size N.

2.1.Empirical insight 2.2.Analytical results

• Verification of the analytical formula $(N = 6 \text{ and } N_{exp} = 1000)$



 \implies Very good agreement between empirical and analytical results.

• Following Daley (1991), it can be shown that the noise length-scale is

$$L_{V^e} = \frac{L_{\epsilon^b}}{\sqrt{2}}.$$

 \implies The sampling noise \mathbf{V}^e is **smaller scale** than the bkg-error field.

3.1.Formulation

Objective filtering of sampling noise (Raynaud et al., 2009)

Notations

Let's ${\mathcal S}$ be the spherical spectral transform, we define the spectral fields :

$$\tilde{\mathbf{S}} = \mathcal{S}(\tilde{\mathbf{V}}) \quad \tilde{\mathbf{S}}^{\star} = \mathcal{S}(\tilde{\mathbf{V}}^{\star}) \quad \mathbf{S}^{e} = \mathcal{S}(\mathbf{V}^{e})$$

Formulation

An objective filter $\boldsymbol{\rho}$, such that $\tilde{\mathbf{S}}^{\star}(n,m) \sim \boldsymbol{\rho}(n)\tilde{\mathbf{S}}(n,m)$, is defined by (Berre et al. (2007))

$$\rho = \frac{1}{1 + \frac{\mathbf{P}(\mathbf{S}^e)}{\mathbf{P}(\tilde{\mathbf{S}}^{\star})}}$$
, where $\mathbf{P}(.)$ is the power spectrum.

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• ρ is a simple function of the **noise/signal ratio**,

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$$\rho = \frac{1}{1 + \frac{\mathbf{P}(\mathbf{S}^e)}{\mathbf{P}(\tilde{\mathbf{S}}^{\star})}}, \text{ where } \mathbf{P}(.) \text{ is the power spectrum.}$$

- ρ is a simple function of the **noise/signal ratio**,
- it can be estimated with the help of the $E[\mathbf{V}^{e}\mathbf{V}^{e^{T}}]$ formula.

4.1.Experimental setup 4.2.Results

Application to the Arpège model σ^{b} (Raynaud et al., 2009)

Experimental setup

- Météo-France Arpège operational model
- Ensemble of 6 independent 3D-Fgat assimilation experiments (Berre et al., 2007, operational since July 2008) :
 - explicit perturbation of observations
 - implicit perturbation of background
 - perfect model framework











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4.1.Experimental setup 4.2.Results



VO surface

VO 500hPa

5.1.Impact of the filtering procedure 5.2.Impact of errors "of the day"

Does spatial filtering of variances have an impact in the (very) end? (Raynaud et al., 2009)



2 impact studies from 15/02/08 to 20/03/08, using :

- vorticity variances "of the day",
 - either raw
 - or objectively filtered
- climatological variances for other variables.

5.1.Impact of the filtering procedure

5.2.Impact of errors "of the day"



(b) 6h-forecast of 850 hPa geopotential over Euratl

So, the response is **YES**! Spatial filtering has a positive impact.

5.1.Impact of the filtering procedure 5.2.Impact of errors "of the day"

Impact of errors "of the day" on an extreme weather event : case of the french storm of 10 February 2009



- 48h-forecasts using :
 - climatological variances
 - variances "of the day" (including VO,D,T,Ps,Q)

- Analysis valid on 10/02/09 at 00 UTC

About the filtering of variances

- Close link between spatial structures of background-error and sampling noise
- Objective filter based on noise-to-signal ratio
- In a NWP context, this filter is robust and nearly cost-free
- Filtered variance maps accurately reflect the underlying flow

About the filtering of variances

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About impact

- Filtered variances improve the background fit to observations and provide more accurate forecasts than raw variances
- The use of a complete set of variances "of the day" results in better forecasts, especially in cases of intense weather events

Perspectives

- Validation and tuning of the filtered variances (Desroziers et al., 2005).
- \bullet Use of such filtered flow-dependent variances in the operational Arpège ${\bf B}$ matrix.
- Ultimate goal : combined use of filtered flow-dependent variances and correlations (*Pannekoucke et al., 2007*).

Perspectives

- Validation and tuning of the filtered variances (Desroziers et al., 2005).
- Use of such filtered flow-dependent variances in the operational Arpège **B** matrix.
- Ultimate goal : combined use of filtered flow-dependent variances and correlations (*Pannekoucke et al., 2007*).

Thank you for your attention!

• Estimation of the noise spectrum $\mathbf{P}(\mathbf{S}^e)$ $E[\mathbf{V}^e \mathbf{V}^{e^T}] = \frac{2}{N-1} \tilde{\mathbf{B}}^{\star} \circ \tilde{\mathbf{B}}^{\star} \rightarrow \text{needs to estimate } \tilde{\mathbf{B}}^{\star} = E[\tilde{\mathbf{B}}]$

With *Ergodic + homogeneous* hypotheses :

$$E[\tilde{\mathbf{B}}_{j.}] \approx \frac{1}{N_t N_i} \sum_{t=1,i=1}^{N_t,N_i} \tilde{\mathbf{B}}_{i.}(t) ,$$

- B̃_i is the local spatial covariance at gridpoint *i*,
- N_t is the number of dates in the time average,
- N_i is the number of gridpoints over the globe.

In the isotropic case, $E[\tilde{\mathbf{B}}_{j.}] = \overline{C}$ (1D) and :

$$\mathbf{P}(\mathbf{S}^e) = \frac{2}{N-1} L(\overline{C}^2)$$

 \bullet Estimation of the noise-free variance spectrum $\mathbf{P}(\tilde{\mathbf{S}}^{\star})$

$$\mathbf{P}(\tilde{\mathbf{S}}^{\star}) = \mathbf{P}(\tilde{\mathbf{S}}) - \mathbf{P}(\mathbf{S}^{e})$$